## Supersonic flow of a Bose-Einstein condensate past an oscillating attractive-repulsive obstacle

E. G. Khamis<sup>1,2,\*</sup> and A. Gammal<sup>1,†</sup>

<sup>1</sup>Instituto de Física, University of São Paulo, 05508-090 São Paulo, Brazil <sup>2</sup>Department of Mathematical Sciences, Loughborough University, Loughborough, LE11 3TU, United Kingdom (Received 21 December 2012; published 1 April 2013)

We investigate by numerical simulations the pattern formation after an oscillating attractive-repulsive obstacle inserted into the flow of a Bose-Einstein condensate. For slow oscillations we observe a complex emission of vortex dipoles. For moderate oscillations organized lined up vortex dipoles are emitted. For high frequencies no dipoles are observed but only lined up dark fragments. The results shows that the drag force turns negative for sufficiently high frequency. We also successfully model the *ship waves* in front of the obstacle. In the limit of very fast oscillations all the excitations of the system tend to vanish.

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Introduction. The realization of Bose-Einstein condensate (BEC) in atomic gases has boosted intense theoretical and experimental investigation about its exotic properties. BEC is a paradigm of a quantum fluid and in a weak interacting case can well be described by the Gross-Pitaevskii equation [1]. Eventually BEC spread to other systems such as exciton polaritons [2], offering new possibilities for experimental tests. An interesting feature of a quantum fluid is its contrasting behavior as compared with a classical fluid. The flow of a quantum fluid past an obstacle only generates drag force above a certain subsonic critical velocity and the energy can be dissipated into collective excitations of the fluid. This dissipation can be inferred from numerical experiments by the mean drag on the obstacle [3]. A superfluid behavior is revealed below this velocity where nucleation of vortices never occurs and no excitations are generated [4].

There is also a supersonic critical velocity where oblique vortex streets are transformed into stable oblique dark solitons [5]. For higher velocities, the general picture of the diffraction pattern in the supersonic flow past a disk-shaped impenetrable obstacle consists of two different parts separated by the Mach (or Cherenkov) cone [6]. Outside the Mach cone there is a region of linear waves that we will refer to simply as ship waves [7,8]. Inside the Mach cone a pair of oblique dark solitons is gradually formed behind the obstacle if the radius of the obstacle is of healing length order. For greater radius more pairs of oblique solitons are generated. Interaction of solitons was studied in [9,10] where it was found that the angle between dark solitons decreases as the obstacle radius increases for a fixed supersonic velocity of the flow. In previous experimental works [7,11,12] the existence of such nonlinear structures were suggested. However, only recently the generation of stable oblique dark solitons was experimentally demonstrated in the flow of a Bose-Einstein condensate of exciton polaritons past an obstacle [13,14]. A numerical study to support these experimental findings was done in [15], and the observation of vortex dipoles in an oblate atomic Bose-Einstein condensate [16] suggests that the supersonic studies can also be carried in this system.

In atomic BEC, obstacles are typically represented by detuned lasers that can effectively be attractive (red detuned) or repulsive (blue detuned) obstacles, by the use of Feshbach resonances. The first numerical study of attractive obstacles was done in Ref. [17] where it was established the critical velocity to the formation of vortices and corrects the velocity found in [3] in the case of repulsive obstacles. Numerical studies revealed that turbulence can also be achieved and studied by spatial oscillation of a repulsive obstacle [18]. A clever way to control the formation of vortices moving attractive and repulsive laser beams was proposed in [19]. The disappearance of gray soliton and phonon excitations was demonstrated in [20] by oscillating a repulsive obstacle in a quasi-one-dimensional trapped BEC at high obstacle velocities. It was found in [21] that vibration of an obstacle modulates the vortex street.

In the present work, we study the flow of a BEC past an oscillating attractive and repulsive obstacle. The motivation is to answer the question: *Can we get rid of drag for very fast oscillations?* We investigate different regimes from slow to very fast oscillations. Since we are working in the supersonic regime we can divide the study inside and outside the Mach cone as follows.

*Model equations.* We consider the flow of an atomic BEC past an obstacle in the framework of the Gross-Pitaevskii (GP) mean-field approach. In the rest frame, the condensate is well described by the macroscopic wave function  $\Psi \equiv \Psi(x, y, z, t)$  obeying the time-dependent GP equation

$$i\hbar\frac{\partial\Psi}{\partial t} = -\frac{\hbar^2}{2m}\nabla^2\Psi + U_{\rm ext}\Psi + \frac{4\pi a\hbar^2}{m}|\Psi|^2\Psi, \quad (1)$$

where  $\nabla^2 \equiv \partial_x^2 + \partial_y^2 + \partial_z^2$ , the external potential  $U_{\text{ext}} = U_{\text{trap}}(x, y, z) + U(x + vt, y, z, \Omega t)$  is represented by the sum of a harmonic trap  $U_{\text{trap}}$  and a time-dependent obstacle potential U that oscillates with frequency  $\Omega$ , m is the atomic mass, and a is the *s*-wave scattering length.

We will limit our study to the case of the quasi-twodimensional (2D) limit, i.e., we have a strong harmonic confinement in the *z* direction. In this regime we can approximate  $\Psi(x,y,z,t) = \psi(x,y,t)\phi(z)e^{-i\mu_z/\hbar}$ , where  $\phi(z)$  and  $\mu_z$  are the ground state and energy, respectively, for the confinement in the *z* direction [19,22]. Substituting in Eq. (1) and integrating

<sup>\*</sup>e.g.khamis@lboro.ac.uk

<sup>†</sup>gammal@if.usp.br

in the z direction we obtain

$$i\hbar\frac{\partial\psi}{\partial t} = -\frac{\hbar^2}{2m}\left(\partial_x^2 + \partial_y^2\right)\psi + U\psi + g|\psi|^2\psi, \quad (2)$$

where  $g = 4\pi a\hbar^2 m^{-1} \int \phi^4(z) dz$  is the effective interaction in two dimensions. We consider here that the obstacle runs close to the center of the trap. In this region the condensate is almost homogeneous and the potential in the x and y directions is weak as compared to the obstacle potential. So the harmonic potential is neglected for studying the excitation caused by the obstacle.

We introduce dimensionless variables  $\tilde{x} = x/\xi$ ,  $\tilde{y} = y/\xi$ ,  $\tilde{t} = gn_0t/\hbar$ ,  $\tilde{\psi} = \psi/\sqrt{n_0}$ ,  $\tilde{U} = U/gn_0$ , and  $\tilde{\Omega} = \Omega\hbar/gn_0$ , the Mach velocity  $M = v/c_s$ , where  $n_0$  is a characteristic 2D density of atoms at the center of the trap,  $\xi = \hbar/\sqrt{mn_0g}$  is the characteristic length, and the sound velocity  $c_s = \hbar/m\xi$ . Typical experimental values are  $\xi \sim 0.3 \ \mu\text{m}$  and  $\hbar/gn_0 \sim$ 0.18 ms [16]. Thus for  $\tilde{\Omega} \sim 1$  we have oscillations of the order of kHz well within experimental reach. Substituting in Eq. (2) and after dropping the tildes for convenience we get

$$i\frac{\partial\psi}{\partial t} = -\frac{1}{2}\left(\partial_x^2 + \partial_y^2\right)\psi + U\psi + |\psi|^2\psi, \qquad (3)$$

where  $U = U(x + Mt, y, \Omega t)$ .

For computational purposes, in Eq. (3) we make a global phase transformation  $\psi' = e^{it}\psi$  and later a Galilean transformation x' = x + Mt, t' = t leading to

$$i\frac{\partial\psi}{\partial t} = -\frac{1}{2}\left(\partial_x^2 + \partial_y^2\right)\psi - iM\partial_x\psi - \psi + |\psi|^2\psi + U\psi,$$
(4)

where  $U = U(x, y, \Omega t)$ , the primes were omitted for convenience, and subscripts here mean derivatives. This equation describes the system in the obstacle reference frame. The obstacle is a laser beam that continuously oscillates from blue detuned to red detuned and vice versa, which can be written as

$$U(x, y, \Omega t) = U_0 \cos(\Omega t) \exp\left[\frac{-2(x^2 + y^2)}{w_0^2}\right],$$
 (5)

where  $U_0$  and  $w_0$  are the amplitude and the beam waist of the laser, respectively, and  $\Omega = 2\pi/T$  is the oscillation frequency of the detuning in a period T.

Inside the Mach cone. We have solved Eq. (3) numerically in the supersonic regime using  $U_0 = 25$  and  $w_0 = 1$ . In Fig. 1 we show the results for supersonic flow for different oscillation frequencies.



FIG. 1. (Color online) Density diffraction pattern at fixed time t = 100 with the set of parameters: Mach number M = 3, potential intensity  $U_0 = 25$  (the laser beam starts repulsive), and width  $w_0 = 1$  for frequencies  $\Omega = 0, 0.5, 1.5, 10, \text{ and } 20$ .



FIG. 2. (Color online) Left panel: Phase of the diffraction pattern that corresponds to Fig. 1 for frequency  $\Omega = 1.5$  which provides us evidence of vortex dipole formation. Right panel: inset showing vector velocities of the vortex dipole in the selected region. Note that the velocity in the *x* direction was not shifted by the constant flow velocity *M* for visualization purposes.

with

In the case of  $\Omega = 0$  we reproduce the formation of oblique dark solitons [5,6,10]. Outside the Mach cone there is a stationary wave pattern created by interference of linear waves. Inside the Mach cone there are two oblique dark solitons that decay at the end points into vortices, situated symmetrically with respect to the direction of the flow. As we turn on the oscillations observe the emission of dark fragments. For  $\Omega = 0.5$  these fragments can be identified as vortex dipoles and form a pattern of "5 in a dice." As the frequency is increased to  $\Omega = 1.5$  these fragments stand well aligned as vortex dipoles as can be identified by the phase plot (see Fig. 2). These dipoles are followed by a secondary radiation emission, identified as a straight line almost parallel to the Mach cone. As the frequency is further increased to  $\Omega = 10$  the fragments can no longer be identified either as a single vortex or as vortex dipoles. By looking at the phase the fragments are identified as short gray solitons that propagate obliquely to the flow, analogous to the ones observed in Ref. [23]. To check the (non)vorticity character after some time of fragments formation we turned off the intensity of the obstacle and no decay into vortices was observed.

One can explain the general behavior as follows. For  $\Omega = 1.5$  the oscillation acts as a "chopper" that turns *on* and *off* the



dipole emission. In this specific case the *on* time is more than enough to generate vortex dipoles and thus we have excess of energy that is ejected as secondary radiation. In the fast oscillating regime ( $\Omega = 10$ ) the time the oscillation is *on* is not enough to form dipoles and just small dark solitons can be seen. As the frequency is around  $\Omega = 20$  practically no more fragments can be seen. To check the consistency of our analysis we studied the number of fragments as a function of  $\Omega$ . One can estimate that rate of fragments emission is close to 1, meaning that at each period one fragment is emitted. The linear behavior confirms the modeling of the oscillating obstacle as a chopper.

*Outside the Mach cone.* Ship waves are formed in front of the obstacle. The theory for a nonoscillating obstacle was previously studied for a  $\delta$  function in Ref. [24], where it was found that the density changes are given by

$$\delta n = V_0 q \, \cos\left(\Phi - \pi/4\right)$$

(6)

$$q \equiv \sqrt{\frac{2k}{\pi r}} \frac{[(M^2 - 2)k^2 + 4(M^2 - 1)]^{1/4}}{[(M^2 - 2)k^2 + 6(M^2 - 1)]^{1/2}},$$



FIG. 3. (Color online) Left panel: Coordinates that define the radius-vector **r** and the wave vector **k**. The latter one is normal to the wave front which is shown schematically by a curved line. Right panel: Numerically calculated wave pattern for a fast oscillating obstacle at fixed time t = 100 with the set of parameters M = 2,  $U_0 = 25$ ,  $w_0 = 1$ , and  $\Omega = 40$ . Solid line (red) corresponds to linear analytical theory [Eq. (7)] for the line of constant phase.

 $k \equiv 2\sqrt{M^2 \cos^2 \eta - 1}$ ,  $\Phi \equiv kr \cos \mu$ , and the angles  $\mu$  and  $\eta$  are defined according to Fig. 3.

According to [24], one can find the shape of the lines of constant phase (wave crests)  $\Phi = kr \cos \mu$  in a parametric form

$$x = r \cos \chi = \frac{4\Phi}{k^3} \cos \eta (1 - M^2 \cos 2\eta),$$
  

$$y = r \sin \chi = \frac{4\Phi}{k^3} \sin \eta (2M^2 \cos^2 \eta - 1).$$
(7)

Predictions of the analytical theory are compared with the numerically calculated wave pattern in Fig. 3 and excellent agreement is found. So, the theory previously developed in [24] remains valid even for a fast oscillating obstacle.

For a fast oscillating we assume that the resulting ship waves can be computed by the Huygens principle, i.e., by the superposition of stationary densities generated by obstacles at different positions along the flow. Averaging over a period this can be expressed as  $\delta n_{\rm osc} =$  $V_0 q(1/T) \int_0^T \cos(\Omega t) \cos(\Phi - \pi/4 + kMt) dt$ , where the term kMt was added representing phase change due to the obstacle movement along the time. After integration in time one obtains

$$\delta n_{\rm osc} = \frac{V_0 q}{2\pi} \left( \frac{-\Omega k M}{\Omega^2 - k^2 M^2} \right) \\ \times \left[ \sin(\Phi - \pi/4 + k M T) - \sin(\Phi - \pi/4) \right]. \tag{8}$$

This formula shows that with increasing  $\Omega$  the magnitude of the  $\delta n_{osc}$  decreases as  $1/\Omega$  and in the limit of  $\Omega \to \infty$ the ship waves vanish. The plot illustrating this behavior is shown in Fig. 4. As we see, Eq. (8) is accurate enough almost everywhere, for  $\Omega \ge 40$ , except in the small vicinity of the obstacle.

Although the obstacle in the theory is represented by a delta function, our numerical simulations using a narrow Gaussian potential as the obstacle provide results in very good agreement with our extended theory.

Drag force. We computed the drag force in the x direction as  $F_x(t) = \int_{\mathcal{A}} dx \, dy \, |\psi|^2 \, \frac{\partial U}{\partial x}$ , where  $\mathcal{A}$  defines an infinite region



FIG. 5. (Color online) Upper panel: Average drag force in the *x* direction,  $\langle F_x \rangle$ , as a function of the frequency  $\Omega$  for different potential intensities. Middle panel: response function amplitude  $A_1$ as a function of  $\Omega$ . Lower panel: relative phase  $\delta_1$  as a function of  $\Omega$ . Here we can see that the sign of  $\langle F_x \rangle$  depends on the relative phase between the forcing potential and the response main mode.

of the fluid around the obstacle. For practical purposes we took the integration along our whole grid. In Fig. 5 we show the average drag taken at one period of oscillation. For slow oscillation frequency  $\Omega \leq 5$  we observe that drag is decreasing and positive as expected since both nonoscillating attractive and repulsive potentials cause positive drag [17]. However, for  $\Omega \geq 5$  the drag is always negative and vanishes in the limit of  $\Omega \rightarrow \infty$ . So, the answer to the question *can we get rid of drag for very fast oscillations* initially proposed is *yes*. Surprisingly, the mean drag also vanishes at a small region of low frequencies and this is a nonintuitive and remarkable result.

The system can be seen as a forced oscillator. In our case the oscillating potential forces the system and we obtain as output an oscillating  $|\psi|^2$ . Thus the drag force in the *x* direction can



FIG. 4. (Color online) Left panel: Profile of intensity in front of the obstacle for x < 0, y = 0 with the set of parameters M = 2,  $V_0 = 2$ ,  $U_0 = 25$  (the laser beam starts repulsive), and  $w_0 = 1$ . Solid lines (black) correspond to linear analytical theory, Eq. (8), and dashed lines (red) to numerical solution of Eq. (3). Right panel: The solid line corresponds to the higher magnitude value of the  $\delta n_{osc}$  close to x = -40 and the red circles correspond to this magnitude calculated numerically at the same position.

be explicitly written as  $F_x = U_0 \cos(\Omega t) R(t)$ , where R(t) is a response function given by  $R(t) \equiv \int_{\mathcal{A}} (-4x/w_0^2) \exp[-2(x^2 + y^2)/w_0^2] |\psi(x,y,t)|^2 dx dy$ . We observed numerically that R(t)is periodic with period T and thus can be written as a Fourier series as  $R(t) = A_0 + A_1 \cos(\Omega t + \delta_1) + A_2 \cos(2\Omega t + \delta_2) + \cdots$ , where A's are amplitudes and  $\delta$ 's are relative phases to the forcing potential. Averaging the drag force in time we have  $\langle F_x \rangle = (1/T) \int_0^T U_0 \cos(\Omega t) R(t) dt$ , and only the second term of the series survives giving  $\langle F_x \rangle = (U_0 A_1 / \Omega) \cos(\delta_1)$ , thus the sign of  $\langle F_x \rangle$  depends on the relative phase between the forcing potential and the response main mode. We computed  $I \equiv$  $(1/T) \int_0^T U_0 \sin(\Omega t) R(t) dt = -(U_0 A_1 / \Omega) \sin(\delta_1)$ , and from  $\langle F_x \rangle$  and I we obtained  $A_1$  and  $\delta_1$ .

A negative drag can be interpreted as a force in the upwind direction, meaning *propels* the laser. The question of the energy balance can be explained from the oscillating laser. As it attracts and repels the condensate it pumps energy to the system that causes an upwind force to supersede the downwind force due to the movement of the laser. This upwind force is only generated in the moving and oscillating obstacle. For standing (M = 0) oscillating obstacle the system is radially symmetric and no drag is generated.

*Conclusions.* We have studied the wave pattern generated by an oscillating obstacle in the supersonic flow of a quantum fluid. Turning on oscillations causes disruption of the oblique solitons into dipoles. For  $\Omega = 1.5$  the dipoles are emitted organized as a vortex dipole street. For increasing frequencies dipoles change gradually orientation in the clockwise direction and their bunch resembles the oblique solitons. Finally for very high frequencies the angle of emission increases and vortices vanish. For the waves in front of the fast oscillating obstacle, we could further extend the model previously developed for a nonoscillating obstacle. These waves were shown to gradually diminish according to the averaging of emission of linear waves out of phase. Combined results, both ship waves and solitons tend to vanish for high frequencies leading to a vanishing drag. Remarkably, the mean drag also vanishes at a small region of low frequencies during his change of sign from positive to negative. So, even a powerful laser fast oscillating from red to blue detuning could pass through an atomic BEC without generating vortices or solitons. This result could be experimentally checked with existing setups [16]. Analogous experiments could also be performed with condensates of exciton polaritons [13].

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- F. Dalfovo, S. Giorgini, L. P. Pitaevskii, and S. Stringari, Rev. Mod. Phys. 71, 463 (1999).
- [2] J. Kasprzak, M. Richard, S. Kundermann, A. Baas, P. Jeambrun, J. M. J. Keeling, F. M. Marchetti, M. H. Szymańska, R. André, J. L. Staehli, V. Savona, P. B. Littlewood, B. Deveaud, and Le Si Dang, Nature (London) 443, 409 (2006).
- [3] T. Frisch, Y. Pomeau, and S. Rica, Phys. Rev. Lett. 69, 1644 (1992).
- [4] B. Jackson, J. F. McCann, and C. S. Adams, Phys. Rev. Lett. 80, 3903 (1998); T. Winiecki, J. F. McCann, and C. S. Adams, *ibid.* 82, 5186 (1999); C. Nore, C. Huepe, and M. E. Brachet, *ibid.* 84, 2191 (2000); J. S. Stießberger and W. Zwerger, Phys. Rev. A 62, 061601(R) (2000); C. Huepe and M. E. Brachet, Physica D 140, 126 (2000); A. Aftalion, Q. Du, and Y. Pomeau, Phys. Rev. Lett. 91, 090407 (2003); K. Sasaki, N. Suzuki, and H. Saito, *ibid.* 104, 150404 (2010).
- [5] A. M. Kamchatnov and L. P. Pitaevskii, Phys. Rev. Lett. 100, 160402 (2008).
- [6] G. A. El, A. Gammal, and A. M. Kamchatnov, Phys. Rev. Lett. 97, 180405 (2006).
- [7] I. Carusotto, S. X. Hu, L. A. Collins, and A. Smerzi, Phys. Rev. Lett. 97, 260403 (2006).
- [8] Yu. G. Gladush, G. A. El, A. Gammal, and A. M. Kamchatnov, Phys. Rev. A 75, 033619 (2007).
- [9] E. S. Annibale and A. Gammal, Phys. Lett. A 376, 46 (2011).
- [10] E. G. Khamis and A. Gammal, Phys. Lett. A 376, 2422 (2012).

- [11] A. Amo, J. Lefrère, S. Pigeon, C. Adrados, C. Ciuti, I. Carusotto, R. Houdré, E. Giacobino, and A. Bramati, Nat. Phys. 5, 805 (2009).
- [12] A. Amo, D. Sanvitto, F. P. Laussy, D. Ballarini, E. del Valle, M. D. Martin, A. Lemaître, J. Bloch, D. N. Krizhanovskii, M. S. Skolnick, C. Tejedor, and L. Viña, Nature (London) 457, 291 (2009).
- [13] A. Amo, S. Pigeon, D. Sanvitto, V. G. Sala, R. Hivet, I. Carusotto, F. Pisanello, G. Leménager, R. Houdré, E. Giacobino, C. Ciuti, and A. Bramati, Science 332, 1167 (2011).
- [14] G. Grosso, G. Nardin, F. Morier-Genoud, Y. Léger, and B. Deveaud-Plédran, Phys. Rev. Lett. 107, 245301 (2011).
- [15] S. Pigeon, I. Carusotto, and C. Ciuti, Phys. Rev. B 83, 144513 (2011).
- [16] T. W. Neely, E. C. Samson, A. S. Bradley, M. J. Davis, and B. P. Anderson, Phys. Rev. Lett. **104**, 160401 (2010).
- [17] B. Jackson, J. F. McCann, and C. S. Adams, Phys. Rev. A 61, 051603(R) (2000).
- [18] K. Fujimoto and M. Tsubota, Phys. Rev. A 83, 053609 (2011).
- [19] T. Aioi, T. Kadokura, T. Kishimoto, and H. Saito, Phys. Rev. X 1, 021003 (2011).
- [20] A. Radouani, Phys. Rev. A 70, 013602 (2004).
- [21] H. Saito, K. Tazaki, and T. Aioi, arXiv:1211.5869.
- [22] V. M. Pérez-García, H. Michinel, and H. Herrero, Phys. Rev. A 57, 3837 (1998); K. K. Das, *ibid.* 66, 053612 (2002).
- [23] H. Flayac, G. Pavlovic, M. A. Kaliteevski, and I. A. Shelykh, Phys. Rev. B 85, 075312 (2012).
- [24] E. G. Khamis, A. Gammal, G. A. El, Yu. G. Gladush, and A. M. Kamchatnov, Phys. Rev. A 78, 013829 (2008).